7th International Workshop on Ring-Imaging Cherenkov Detectors (RICH 2010) Cassis, France, 3-7 May 2010

The KM3NeT Neutrino Telescope

Uli Katz ECAP / Univ. Erlangen 04.05.2010

ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS

Friedrich-Alexander-Universität Erlangen-Nürnberg

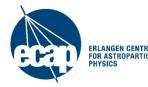






- Technical solutions:
 Decisions and options
- Physics sensitivity
- Cost and implementation
- Summary

KM3NeT



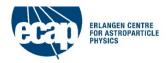
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The Neutrino Telescope World Map



AMANDA

South Pole

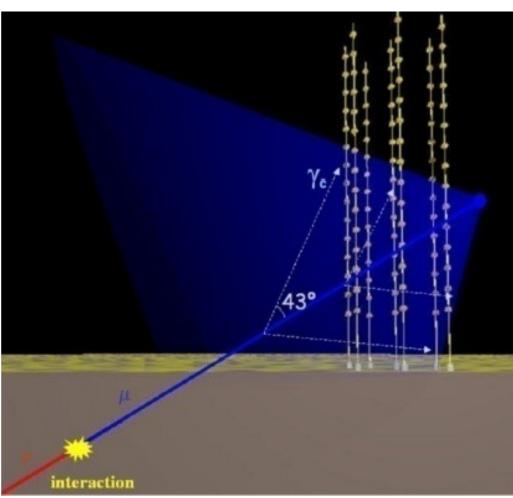


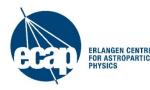
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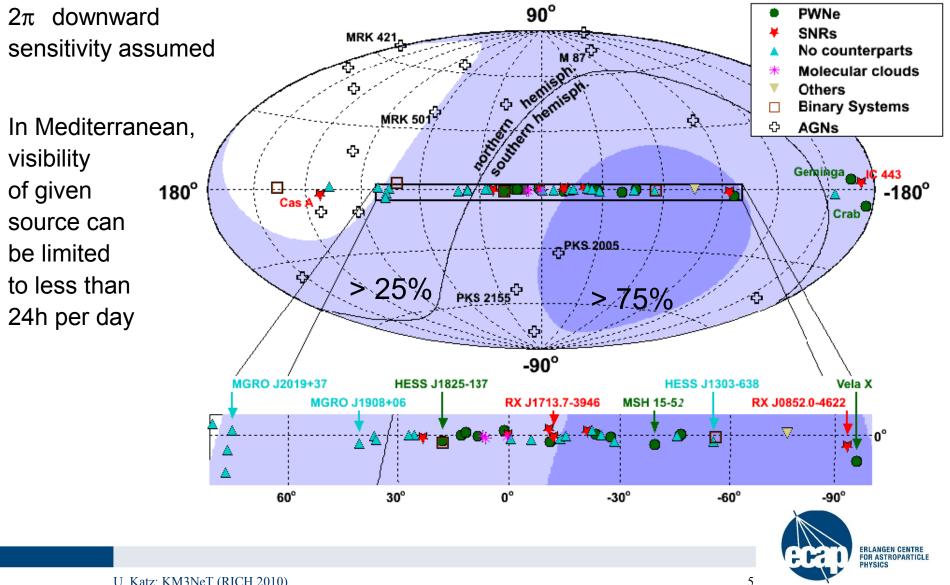
What is KM3NeT ?

- Future cubic-kilometre scale neutrino telescope in the Mediterranean Sea
- Exceeds Northernhemisphere telescopes by factor ~50 in sensitivity
- Exceeds IceCube sensitivity by substantial factor
- Focus of scientific interest: Neutrino astronomy in the energy range 1 to 100 TeV
- Provides node for earth and marine sciences





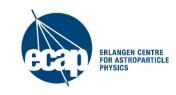
South Pole and Mediterranean Fields of View



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The Objectives

- <u>Central physics goals:</u>
 - Investigate neutrino "point sources" in energy regime 1-100 TeV
 - Complement IceCube field of view
 - Exceed IceCube sensitivity
- Implementation requirements:
 - Construction time ≤5 years
 - Operation over at least 10 years without "major maintenance"



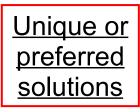
Technical Design

<u>Objective</u>: Support 3D-array of photodetectors and connect them to shore (data, power, slow control)

- Optical Modules
- Front-end electronics & readout
- Readout, data acquisition, data transport
 - Mechanical structures, backbone cable
 - General deployment strategy
- Sea-bed network: cables, junction boxes
 - Calibration devices
 - Shore infrastructure
 - Assembly, transport, logistics
 - Risk analysis and quality control

Design rationale:

Cost-effective Reliable Producible Easy to deploy





Further Challenges

Site characteristics

<u>Objective</u>: Measure site characteristics (optical background, currents, sedimentation, ...)

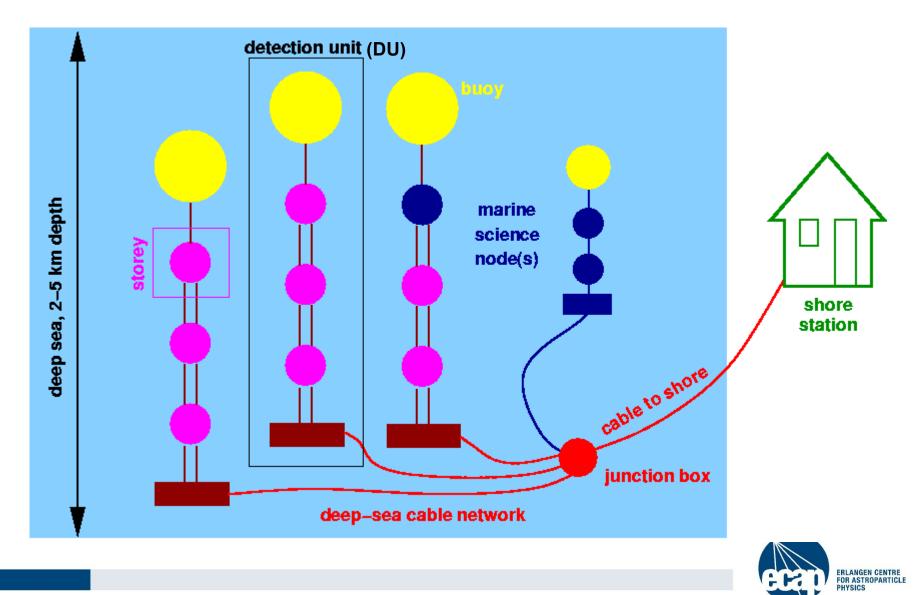
Simulation

<u>Objective</u>: Determine detector sensitivity, optimise detector parameters;

- Earth and marine science node <u>Objective</u>: Design interface to instrumentation for marine biology, geology/geophysics, oceanography, environmental studies, alerts, ...
- Implementation

<u>Objective</u>: Take final decisions (technology and site), secure resources, set up proper management/governance, construct and operate KM3NeT;

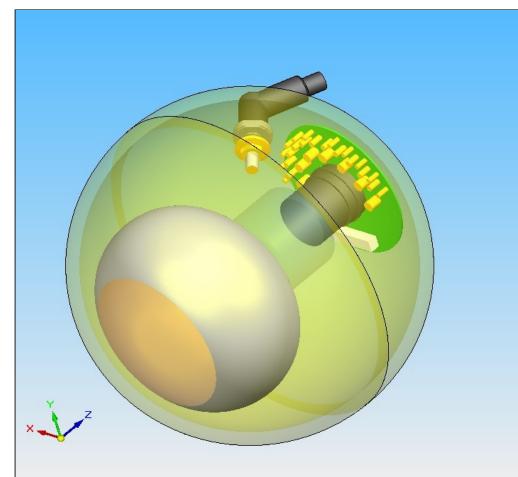
The KM3NeT Research Infrastructure (RI)

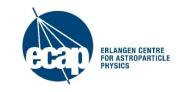


OM "classical": One PMT, no Electronics

Evolution from pilot projects:

- 8-inch PMT, increased quantum efficiency (instead of 10 inch)
- 13-inch glass sphere (instead of 17 inch)
- no valve (requires "vacuum" assembly)
- no mu-metal shielding

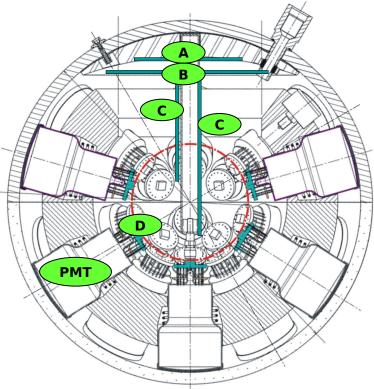




OM with many Small PMTs

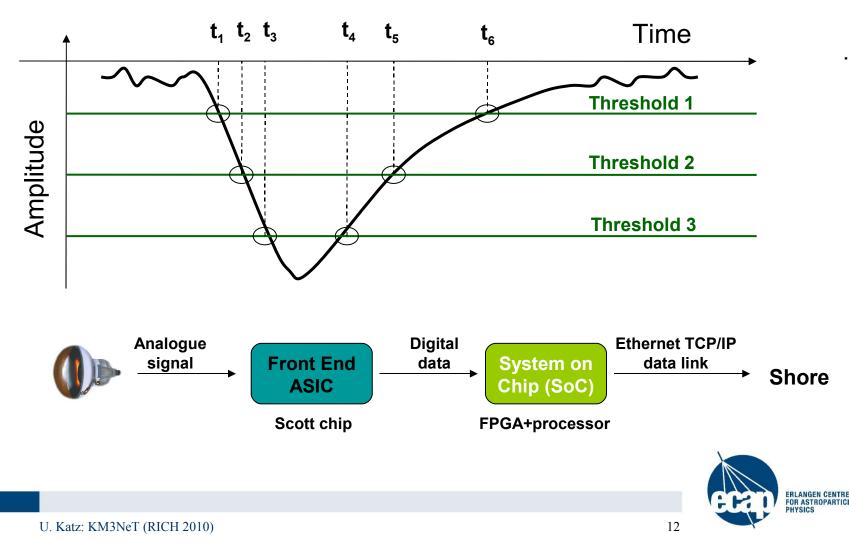
- 31 3-inch PMTs in 17-inch glass sphere (cathode area~ 3x10" PMTs)
 - 19 in lower, 12 in upper hemisphere
 - Suspended by compressible foam core
- 31 PMT bases (total ~140 mW) (D)
- Front-end electronics (B,C)
- Al cooling shield and stem (A)
- Single penetrator
- 2mm optical gel (ANTARES-type)





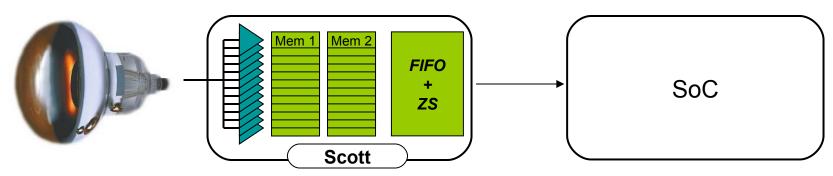
Front-End Electronics: Time-over-Threshold

From the analogue signal to time stamped digital data:

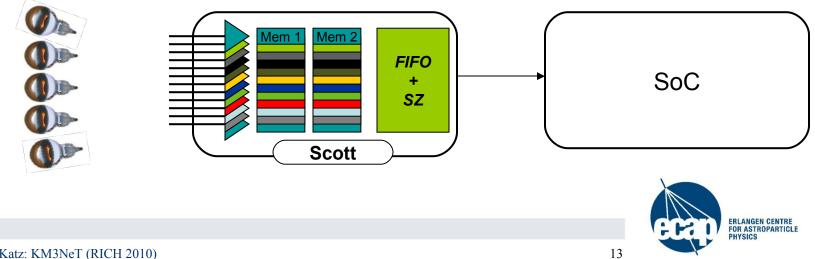


Same Readout for Single- and Multi-PMT OMs

N thresholds for 1 PMT



N/k thresholds for k PMTs



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Data Network

- <u>All data to shore:</u> Full information on each hit satisfying local condition (threshold) sent to shore
- <u>Overall data rate</u> ~ 25 Gbyte/s
- <u>Data transport:</u> Optical point-to-point connection shore-OM Optical network using DWDM and multiplexing Served by lasers on shore Allows also for time calibration of transmission delays
- <u>Deep-sea components</u>: Fibres, modulators, mux/demux, optical amplifiers (all standard and passive)



DUs: Bars, Strings, Triangles

- Flexible towers with horizontal bars
 - Simulation indicates that "local 3D arrangement" of OMs increases sensitivity significantly
 - Single- or multi-PMT OMs
- Slender strings with multi-PMT OMs
 - Reduced cost per DU, similar sensitivity per Euro
- Strings with triangular arrangements of PMTs
 - Evolution of ANTARES concept
 - Single- or multi-PMT OMs
 - "Conservative" fall-back solution

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			er:

Progress in verifying deep-sea technology can be slow and painful

Careful prototype tests are required before taking final decisions

This is a task beyond the Design Study!

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The Flexible Tower with Horizontal Bars

- 20 storeys
- Each storey supports 6 OMs in groups of 2
- Storeys interlinked by tensioning ropes, subsequent storeys orthogonal to each other
- Power and data cables separated from ropes; single backbone cable with breakouts to storeys
- Storey length = 6m
- Distance between storeys = 40 m
- Distance between DU base and first storey = 100m



40m

6m

The Slender String

- Mooring line:
 - Buoy (empty glass spheres, net buoyancy 2250N)
 - Anchor: concrete slab of 1m³
 - 2 Dyneema ropes (4 mm diameter)
 - 20 storeys (one OM each),
 30 m distance, 100m anchor-first storey
- Electro-optical backbone:
 - Flexible hose ~ 6mm diameter
 - Oil-filled

New concept, needs to be tested. Also for flexible tower if successful

One single pressure transition

 Star network between master module and optical modules

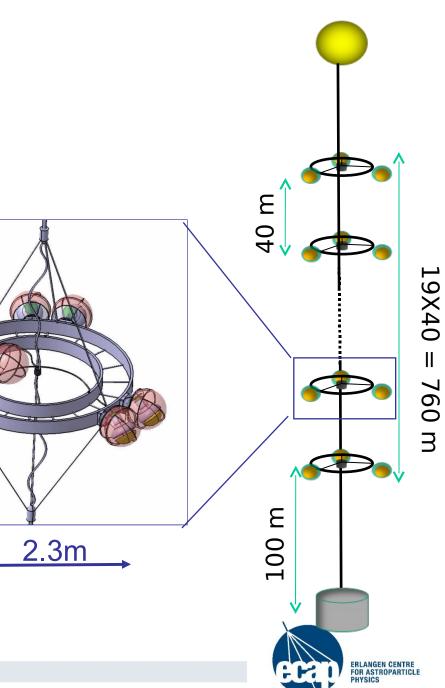




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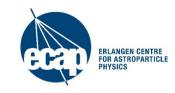
Triangle Structure

- Evolution from ANTARES concept
- 20 storeys/DU, spacing 40m
- Backbone: electrooptical-mechanical cable
- Reduced number of electro-optical penetrations
- Use ANTARES return
 of experience



Deployment Strategy

- All three mechanical solutions: Compact package – deployment – self-unfurling
 - Eases logistics (in particular in case of several assembly lines)
 - Speeds up and eases deployment; several DUs can be deployed in one operation
 - Self-unfurling concepts need to be thoroughly tested and verified
- Connection to seabed network by ROV
- Backup solution: "Traditional" deployment from sea surface



A Flexible Tower Packed for Deployment

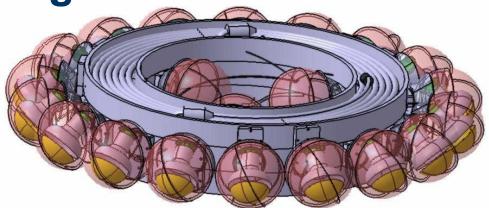




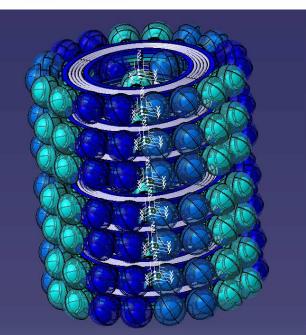
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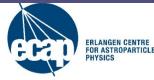
Compactifying Strings

Slender string rolled up for self-unfurling:



3 triangles DU





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Hydrodynamic Stability

- DUs move under drag of sea current
 - Currents of up to 30cm/s observed
 - Mostly homogeneous over detector volume
 - Deviation from vertical at top:

Current	flexible tower	slender string	triangles
[cm/s]	d [m]	d [m]	d [m]
30	84.0	83.0	87.0

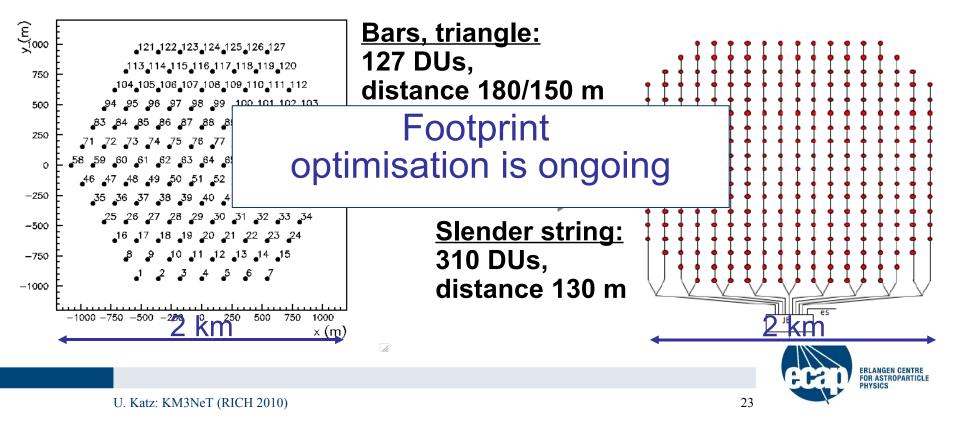
Torsional stability also checked



Detector Building Blocks

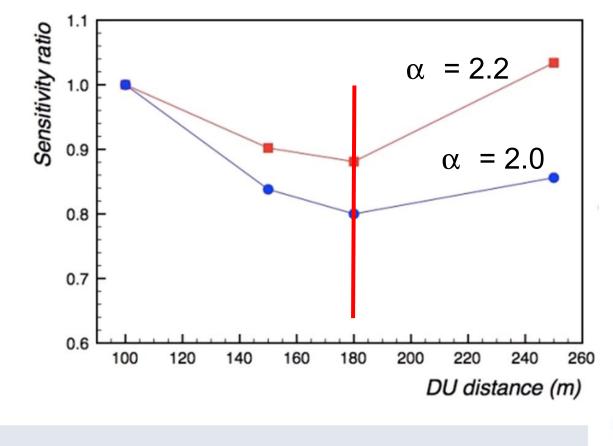
- Different DU designs
 - require different DU distance
 - differ in photocathode area/DU
 - are different in cost

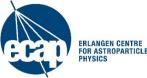




Optimisation Studies

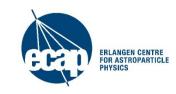
<u>Example</u>: Sensitivity dependence of point-source search on DU distance for flexible towers (for 2 different neutrino fluxes $\sim E^{-\alpha}$, no cut-off)



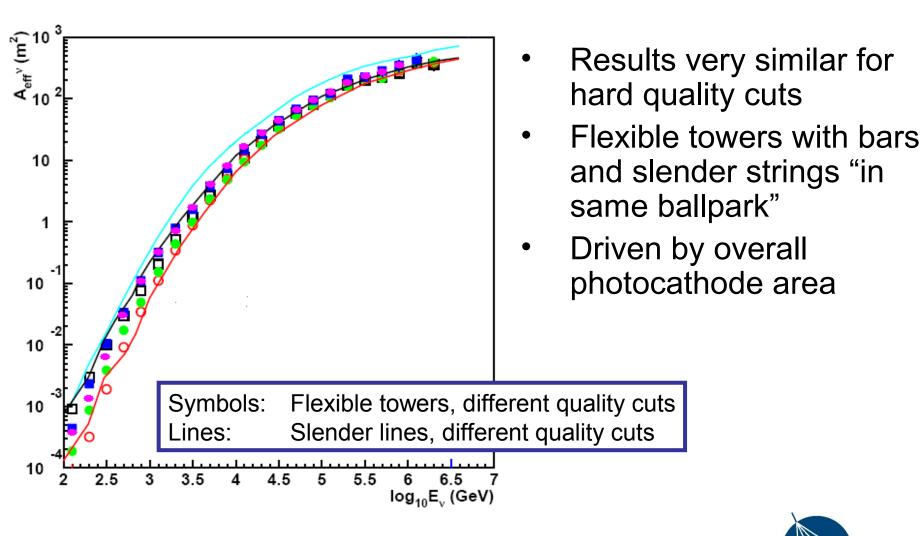


Angular Resolution

median (degree) Investigate \Box distribution of angle between Θ incoming neutrino and reconstructed muon Dominated by 10⁻² [⊥]2 kinematics up to 3 5 6 ~1TeV log₁₀E_v (GeV)



Effective Areas (per Building Block)





Cost Estimates: Assumptions

- Estimate of investment cost
 - no personnel costs included
 - no contingency, no spares
- Assumptions / procedure:
 - Quotations from suppliers are not official and subject to change
 - Common items are quoted with same price
 - Sea Sciences and Shore Station not estimated
 - Estimates worked out independently by expert groups and carefully cross-checked and harmonised thereafter

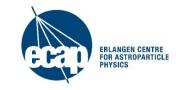


Cost Estimates: Results

• Result of cost estimates (per building block):

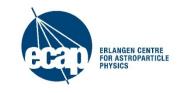
Concept	DU Cost (M€)	No. of DUs	Total DU Cost (M€)	Seafloor Infrastr. (M€)	Deploy- ment (M€)	TOTAL COST (M€)
Flexible towers	0.54	127	68	8	11	87
Slender strings	0.25	310	76	13	14	103
Triangles	0.66	127	83	8	7	99

 Assembly man power (OMs, DU...) is roughly estimated to be 10% of the DU cost

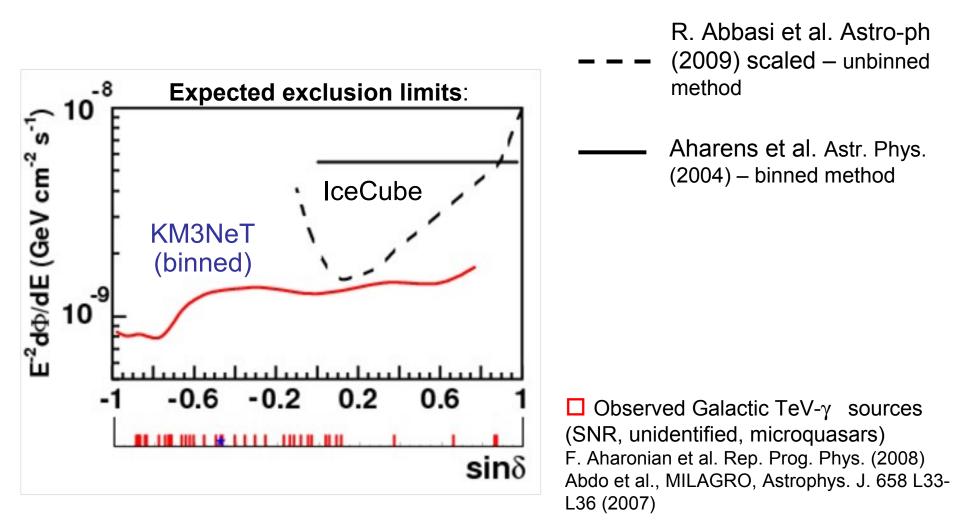


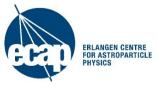
KM3NeT: Full Configuration

- 2 "building blocks" needed to achieve objectives
- Increases sensitivity by a factor 2
- Overall investment ~220 M€
- Staged implementation possible
- Science potential from very early stage of construction on
- Operational costs 4-6 M€ per year (2-3% of capital investment), including electricity, maintenance, computing, data centre and management



Point Source Sensitivity (1 Year)

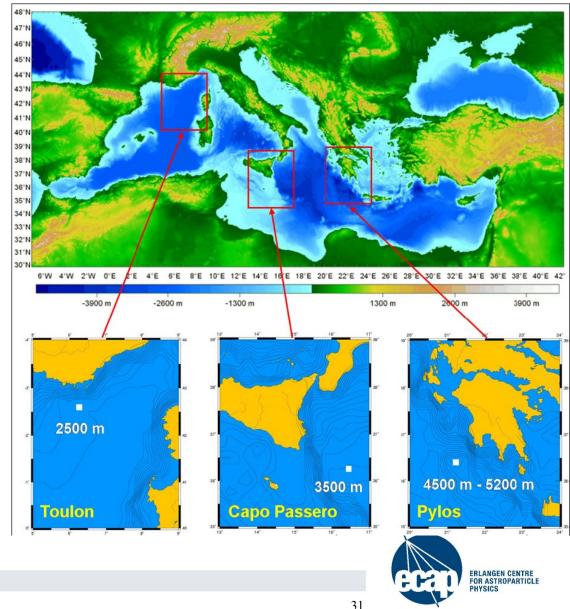




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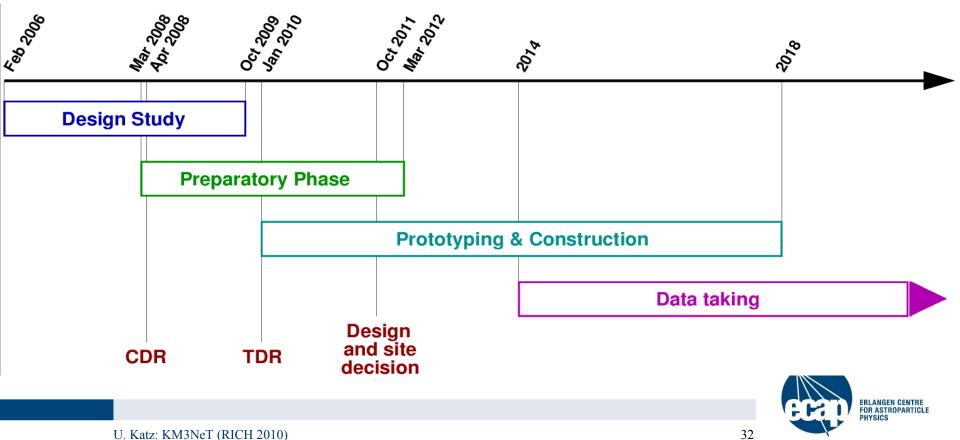
Candidate Sites

- Locations of the three pilot projects:
 - ANTARES: Toulon
 - NEMO: Capo Passero
 - NESTOR: Pylos
- Long-term site characterisation measurements performed
- Site decision requires scientific, technological and political input



Next Steps and Timeline

- Next steps: Prototyping and design decisions
 - TDR public in 3-4 weeks
 - final decisions require site selection
 - expected to be achieved in 18 months
- Timeline:



Conclusions

- A design for the KM3NeT neutrino telescope complementing the IceCube field in its of view and surpassing it in sensitivity by a substantial factor is presented.
- Readiness for construction expected in 18 months
- An overall budget of ~250 M€ will be required. Staged implementation, with increasing discovery potential, is technically possible.
- Within 18 months, remaining design decisions have to be taken and the site question clarified.
- Installation could start in 2013 and data taking soon after.

