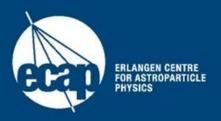
News from KM3NeT

Uli Katz ECAP / Univ. Erlangen 13.10.2013

ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS



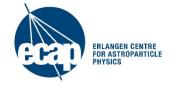


The plan for the next 15 minutes:

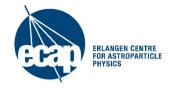
- Decisions taken: Technology and sites
- Tests and prototypes
- Next steps
- Summary

KM3NeT

Not included: Physics aspects & ORCA case study



Decisions taken



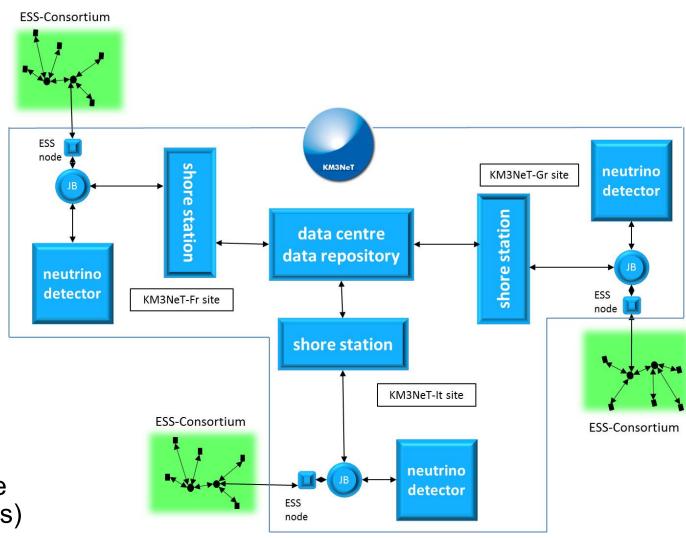
Flashback end-2009 (after Design Study):

- Which architecture to use? (strings vs. towers vs. new designation)
- Design of photo-detection ur (large vs. several small PM sctionality, ...)
- Readout and data acq (how to implement? (how to implement? (how to implement?)
- Deployment tech
 (2 types of "Cc
 ify and unfurl" vs. traditional)
- And finally to) site decision.



KM3NeT: a distributed Research Infrastructure

- Centrally managed
- Common hardware
- Common software, data handling and operation control
- Sites in France, Greece, Italy
- Consistent with funding structure (regional sources)





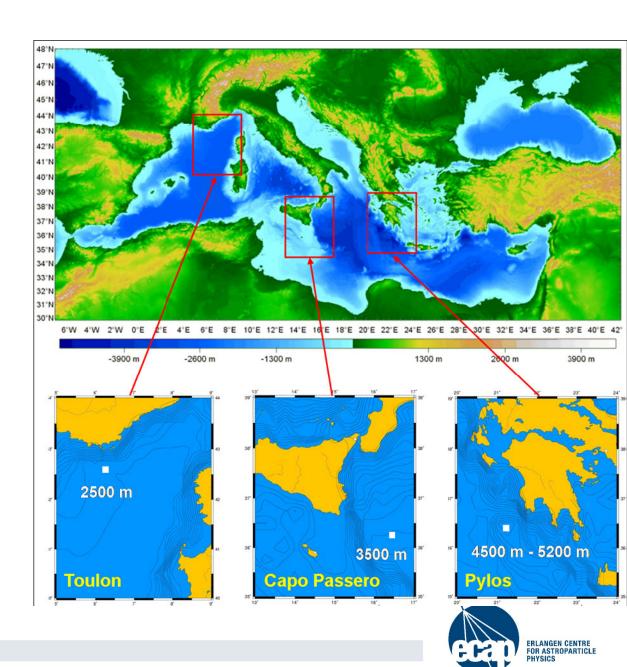
KM3NeT Sites

KM3NeT-France: Toulon

KM3NeT-Italy: Capo Passero

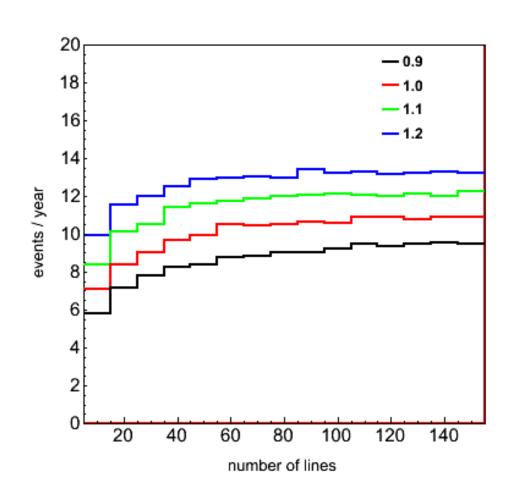
KM3NeT-Greece: Pylos

 Long-term site characterisation measurements performed



The building block concept

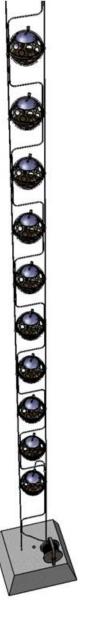
- Building block:
 - 115 detection units
 - Segmentation enforced by technical reasons
 - Sensitivity for muons independent of block size above ~75 strings
 - One block ~ half IceCube
- Geometry parameters optimised for galactic sources (E cut-off)
- Technical feasibility verified
- KM3NeT includes
 6 building blocks





Detection units: Strings

- Mooring line:
 - Buoy (probably syntactic foam)
 - 2 Dyneema[©] ropes (4 mm diameter)
 - 18 storeys (one OM each),
 36m distance, 100m anchor-first storey
- Electro-optical backbone (VEOC):
 - Flexible hose ~ 6mm diameter
 - Oil-filled
 - fibres and copper wires
 - At each storey: connection to 1 fibre+2 wires
 - Break out box with fuses at each storey:
 One single pressure transition



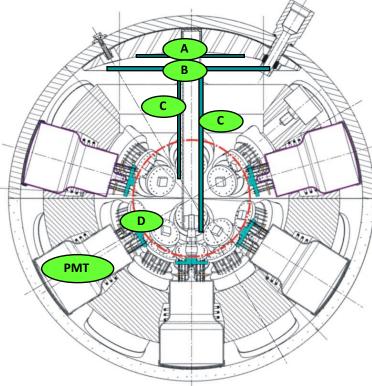




OM with many small PMTs

- 31 3-inch PMTs in 17-inch glass sphere (cathode area~ 3x10" PMTs)
 - 19 in lower, 12 in upper hemisphere
 - Suspended by plastic structure
- 31 PMT bases (total ~140 mW) (D)
- Front-end electronics (B,C)
- Al cooling shield and stem (A)
- Single penetrator
- 2mm optical gel





Advantages of the KM3NeT DOM

- Increased photocathode area
 - 1 KM3NeT DOM = 3 ANTARES OMs
 - Reduces numbers of penetrations/connectors (expensive & risky)
 - Reduces number of optical modules and their infrastructure (expensive)
- 1-vs.-2 photo-electron separation
 - Better sensitivity to coincidences / background suppression
 - Information at online data filter level
- Directionality
 - Additional input to reconstruction and veto algorithms
 - Identification of downgoing events (PMTs are also looking upwards)
 - Reduction of random background (K40, bioluminescence)



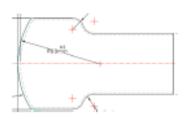
PMT availability

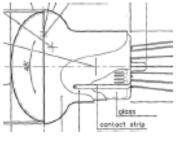
Hamamatsu R12199-02

ET Enterprises Ltd (ETEL, UK)
D783KFLA, D793KFLA – 78 mm
104 pc 10 pc
D792KFLA – 90 mm
12 pc

HZC (China) XP53 7 pc 76 mm

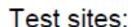




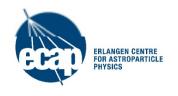






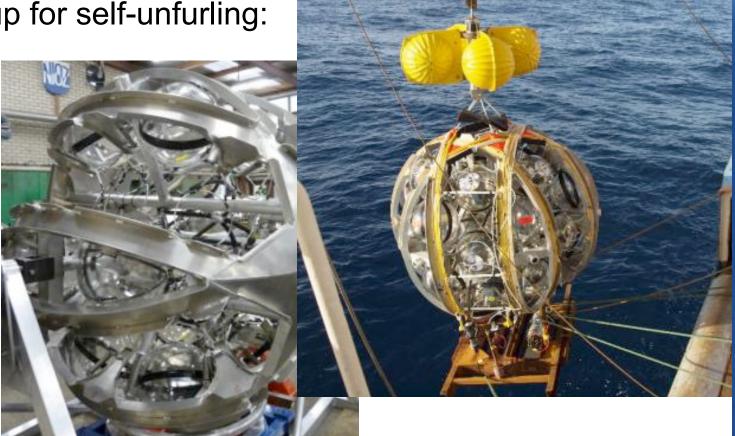


Hellenic University, Nikhef, LNS INFN Catania and ECAP

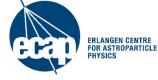


Deploying strings

string rolled up for self-unfurling:







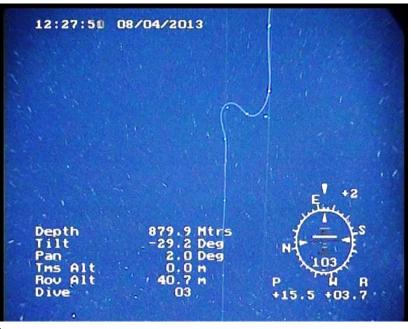
Tests and prototypes



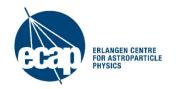
String mechanical deployment tests

Several deployments 2-12 April 2013 at a depth of 1000m (NIOZ boat) 20 miles off the coast of Motril, Spain



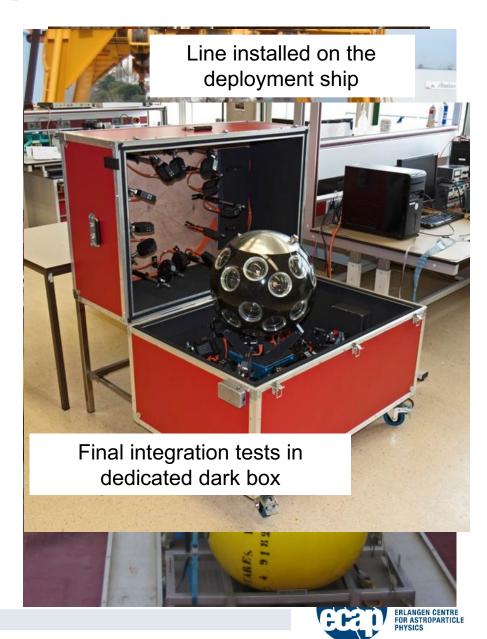


- Successful demonstration of deployment concept
- DOMs are horizontal
- VEOC cable → no leaks
- Some issues with penetrators (understood)
- Second test towards end of year

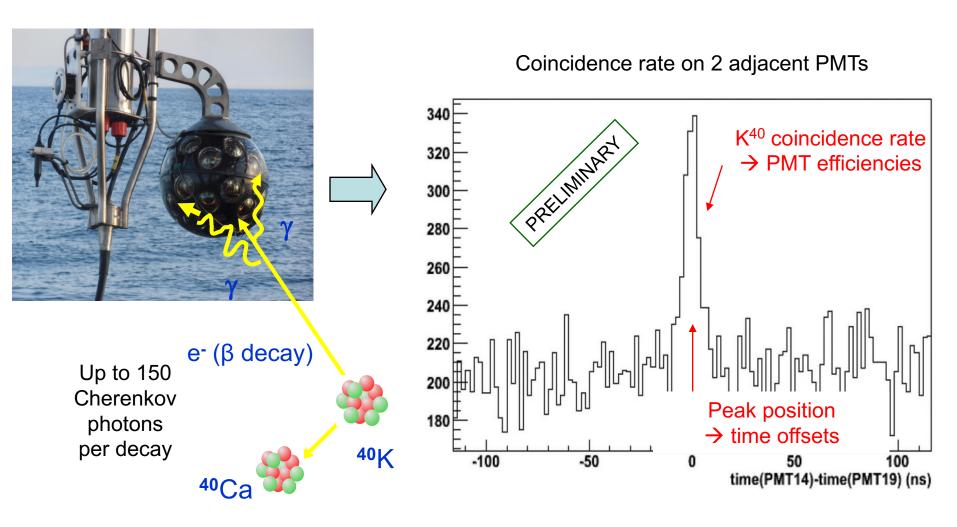


The Pre-production Optical Module

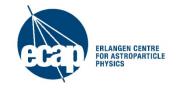
- Fully equipped DOM (31 PMTs + acoustic positioning sensors + time calibration LED beacon)
- Mounted on the Instrumentation Line of ANTARES (2475m deep)
- Internal reference: "PPM-DOM"
- Deployed and connected with ROV on 16 April 2013
- PPM-DOM fully operational and working well



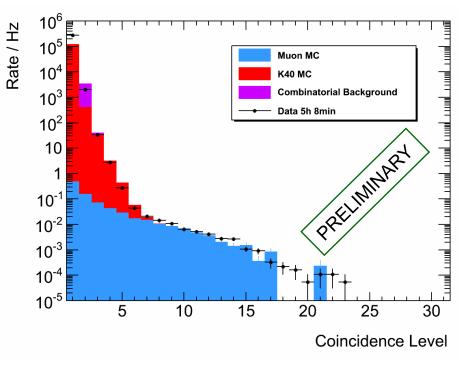
PPM-DOM: K40 Coincidences



Concentration of ⁴⁰K is stable (coincidence rate ~5 Hz on adjacent PMTs)

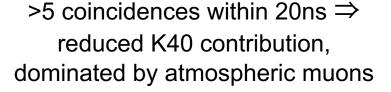


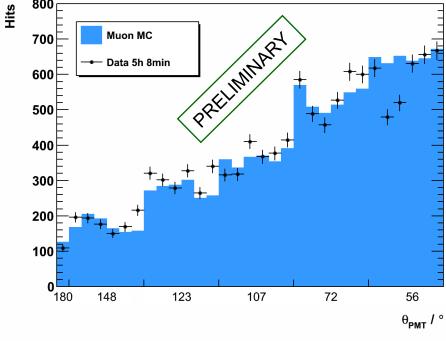
PPM-DOM: Atmospheric Muons



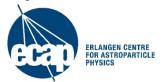
Zenith angle of hit PMTs in events with more than 6 coincident hits



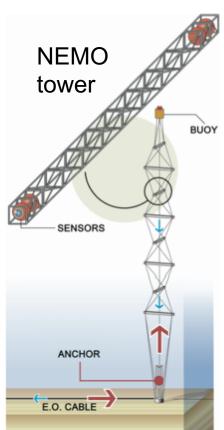




More upper PMTs in multi-hit events ⇒ directional information from single storey

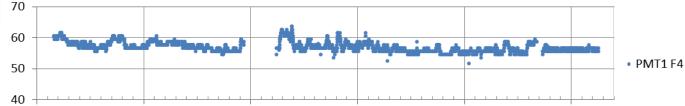


KM3NeT-Italy: site qualification



Connected by ROV March 23, 2013

First continuous rate measurements



 $25/04/2013\ 00:00\ 05/05/2013\ 00:00\ 15/05/2013\ 00:00\ 25/05/2013\ 00:00\ 04/06/2013\ 00:00\ 14/06/2013\ 00:00$

Long term monitoring of site characteristics

Next steps

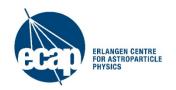


KM3NeT Phase-1

- 40 M€ available (out of ~220 M€ estimated for full KM3NeT)
- Substantial part: European Regional Development Funds Must be spent by end of 2014 → Use or lose!
- KM3NeT decided to embark on first construction phase
 - Transformation consortium → collaboration early 2013, management established, MoU endorsed by collaboration.
 - Construction will start at Toulon and Capo Passero sites, very tight time schedule
 - Common technology, software, data handling, operation, governance

Goals:

- Provide Northern-hemisphere NT with unprecedented sensitivity
- Demonstrate feasibility, operability, stability, resolutions, sensitivity
- Provide infrastructural environment for phase-2



Seabed infrastructure

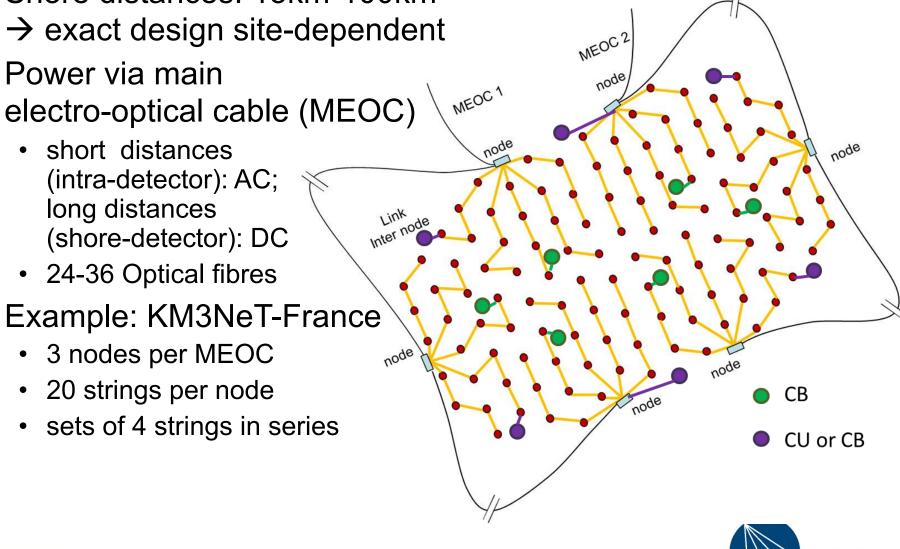
Shore distances: 15km-100km → exact design site-dependent Power via main

> short distances (intra-detector): AC; long distances (shore-detector): DC

• 24-36 Optical fibres

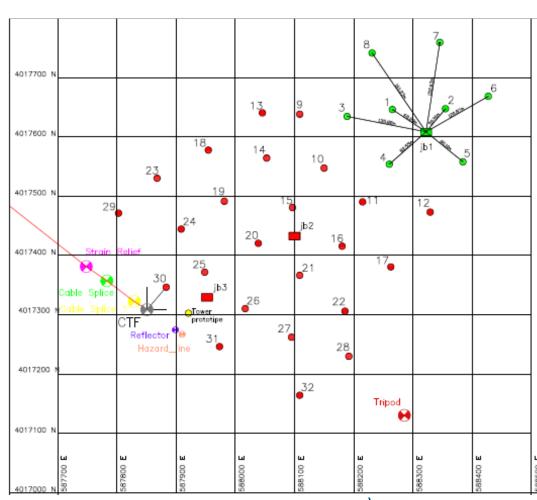
Example: KM3NeT-France

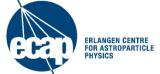
- 3 nodes per MEOC
- 20 strings per node
- sets of 4 strings in series



KM3NeT-Italy installation plans (phase-1)

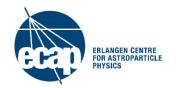
- Start with 8 towers (necessary to match spending profile and to demonstrate construction activity)
- Add 24 strings until 2015
- High-level tower-string data combination





Towards KM3NeT phase-2

- KM3NeT-Greece (beyond phase-1)
 - Application pending
 - If successful: Site development and first detector construction
 - Time scale for decision and implementation unclear
- Full installation (phase-2)
 - No firm commitments yet
 - Financial construction part of phase-1 program;
 EU funding sources (structural funds; Horizon2020; ...)
 - ERIC planned (headquarter in Amsterdam)
- Future of neutrino astronomy will have global dimension

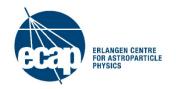


Summary

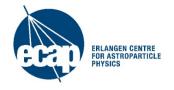


Summary

- KM3NeT will be a distributed, networked research infrastructure.
- Technical design is fixed and decided.
- Intense prototyping and test program ongoing; very encouraging results so far.
- First construction phase will start 2014 (KM3NeT phase-1).
- Path towards full implementation to be defined during phase-1.
- Considering global dimension for future planning will be crucial for neutrino astronomy.

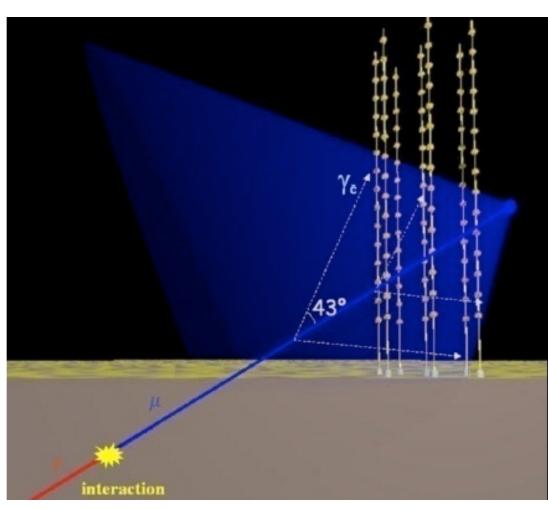


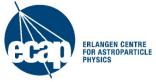
Backup



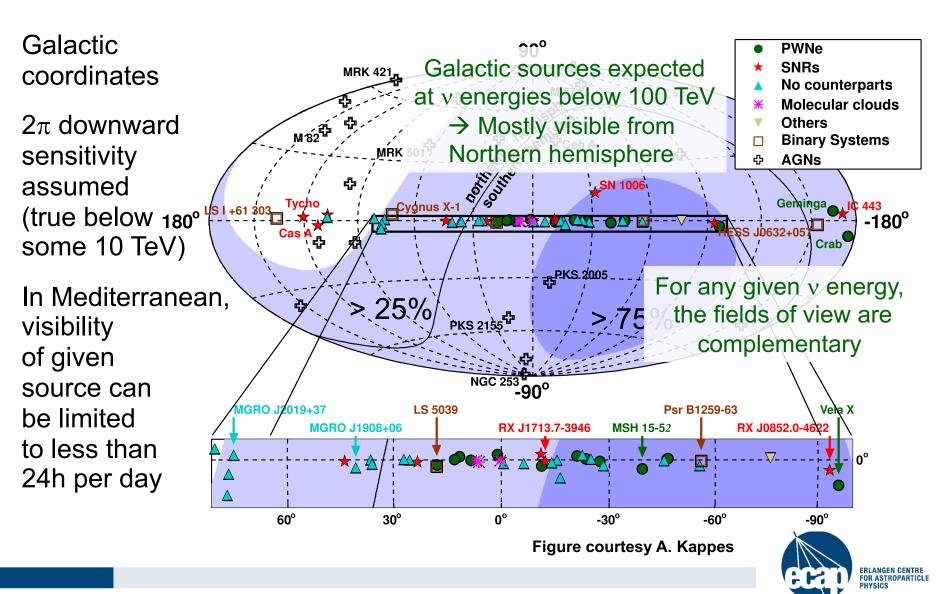
What is KM3NeT?

- Future research infrastructure in the Mediterranean Sea
- Includes cubic-kilometre scale neutrino telescope
- Exceeds Northern-hemisphere telescopes by factor ~50 in sensitivity
- Exceeds IceCube sensitivity by substantial factor
- Provides node for earth and marine sciences





South Pole and Mediterranean fields of view

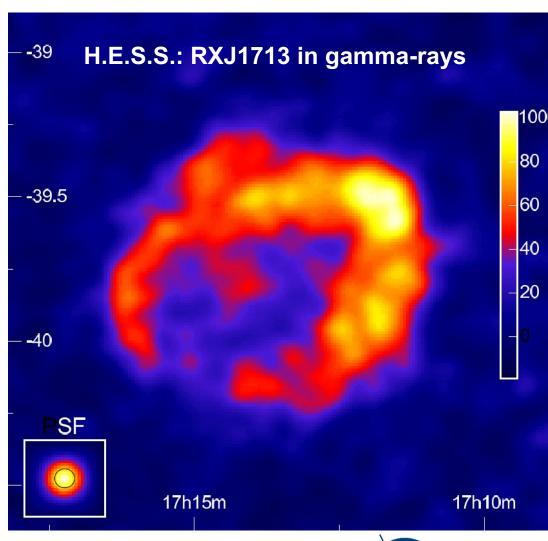


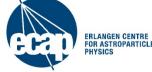
Sensitivity to Galactic sources

- Supernova Remnants
 - RXJ1713.7-3946 (prime example)
 - Vela X
 (exciting option after update of H.E.S. data)

Priority physics objective of KM3NeT

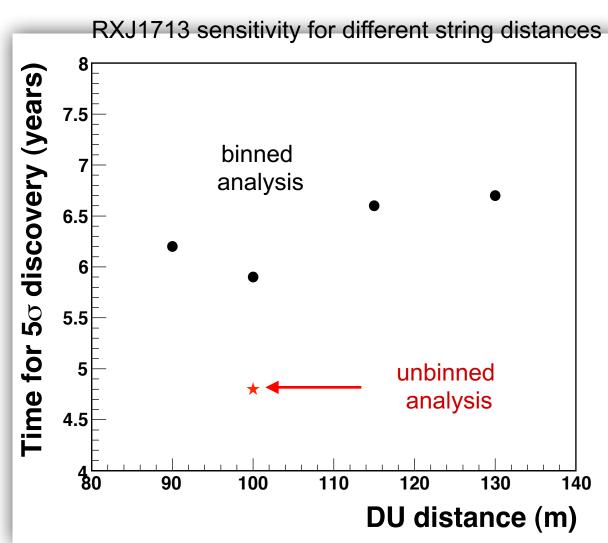
- → Talk Paul Kooijman
- Fermi Bubbles
 - → Talk Paolo Piattelli

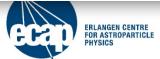




Discovery potential for Supernova remnants

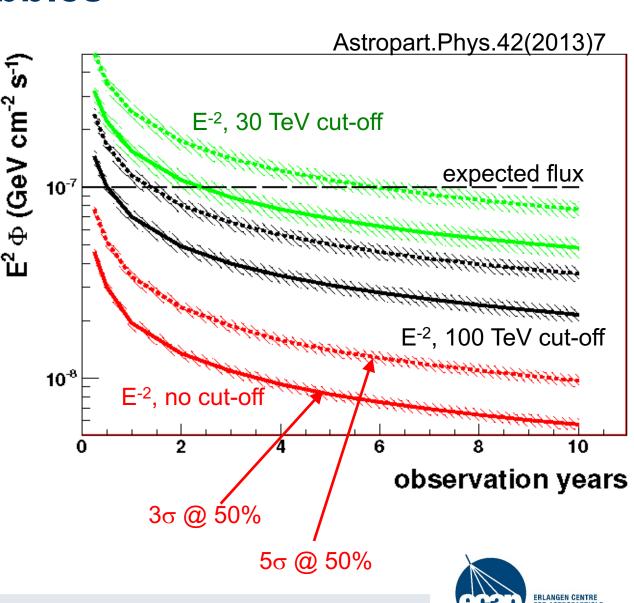
- Simulation results for 2 x 310 strings
- 5σ discovery in less than 5 years for RXJ1713.7-3946 (unbinned analysis)
- Even higher sensitivity for Vela X
- SNR neutrino fluxes
 (E⁻² spectrum with cutoff)
 used for detector
 optimisation





The Fermi bubbles

- Two extended regions above/below centre of Galactic plane
- Fermi detected hard γ emission (E⁻²) up to 100 GeV
- Origin and acceleration has mechanisms under debate if hadronic, hot neutrino source candidate
- Could be first source detected by KM3NeT



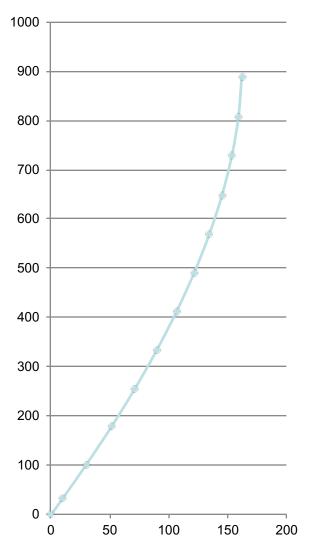
KM3NeT and the new IceCube results

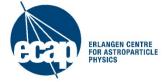
- For Technical Design Report and design optimisation we focused on Galactic sources (µ channel, up-going)
 - Cascade reconstruction and starting track analysis not yet available.
 - In depth-studies under way (high priority).
 - No results ready for this workshop.
- Required: Assumption on the nature of the signal
 - Isotropic (?)
 - Flavour-symmetric
 - ➤ E⁻² flux with cutoff around 2 PeV (?)
- Detector re-optimisation possible for phase-2, not for phase-1

Hydrodynamic stability

- DUs move under drag of sea current
 - Currents of up to 30cm/s observed
 - Mostly homogeneous over detector volume
 - Deviation from vertical at top about 150m at 30cm/s (can be reduced by extra buoyancy)
 - Critical current ~45cm/s (anchor starts to move)

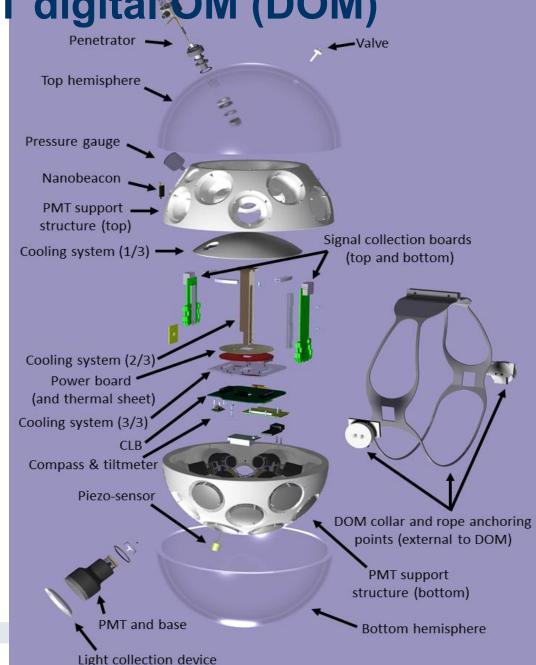
deviation at 30 cm/s





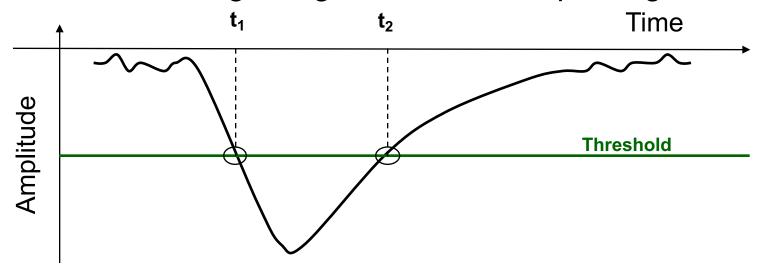
More on the KM3NeT digital OM (DOM)

- Light collection device
 - 20–40% gain in effective photocathode area
- Low power
 - <10 W / DOM</p>
- FPGA readout
 - for each individual PMT
 - sub-ns time stamping
 - time over threshold
- Calibration
 - LED & acoustic piezo
- Optical fibre data transmission
 - DWDM with 80 wavelengths
 - Gb/s readout

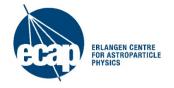


Readout: time-over-threshold

From the analogue signal to time stamped digital data:



- Implemented for each PMT through FPGA on central logic board (CLB) contained in optical module
- All data to shore via optical fibres
- Time synchronisation and slow control
- → see presentations by D. Real and G. Kieft



Deployment strategy

- Compact package deployment self-unfurling
 - Eases logistics (in particular in case of several assembly lines)
 - Speeds up and eases deployment; several units can be deployed in one operation
 - Self-unfurling concept being thoroughly tested and verified
- Connection to seabed network by ROV

