The Neutrino Telescope of the KM3NeT Deep-Sea Research Infrastructure



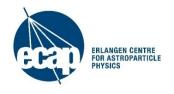
Robert Lahmann for the KM3NeT Consortium Erlangen Centre for Astroparticle Physics TIPP 2011, Chicago 11-June-2011





Outline

- Objectives and Physics Case
- Technical Design and Implementation
- Optical Modules and Electronics

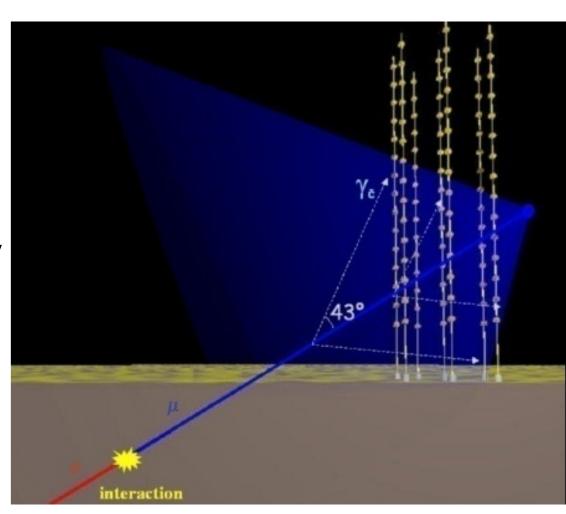


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What is KM3NeT?

- Future cubic-kilometre scale neutrino telescope in the Mediterranean Sea
- Exceeds Northern-hemisphere telescopes by factor ~100 in sensitivity
- Exceeds IceCube sensitivity by substantial factor
- Provides node for earth and marine sciences (continuous deep-sea measurements)





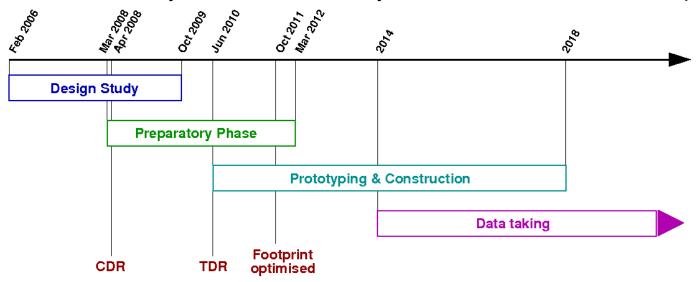
KM3NeT Organisation

The KM3NeT consortium:

40 European institutes
 including those from Antares, Nemo and Nestor neutrino telescope projects

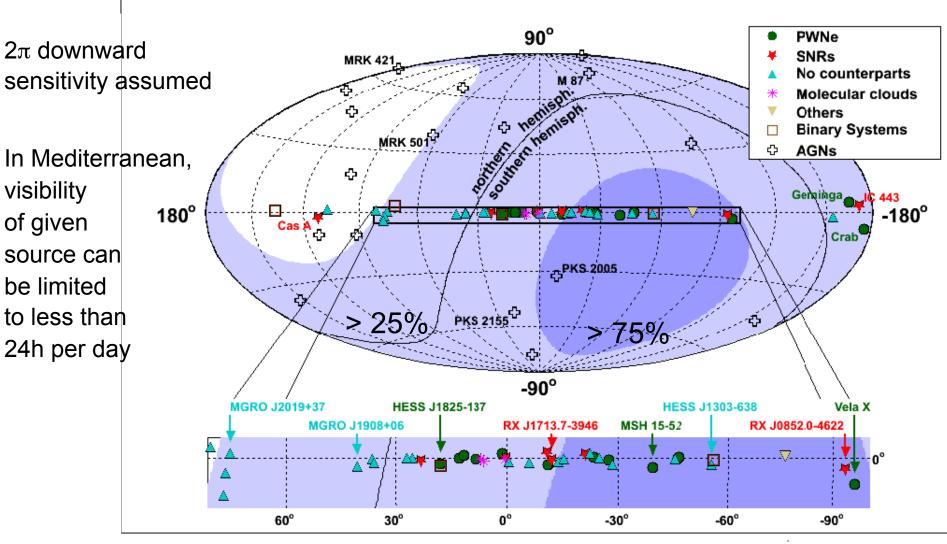
10 countries

(Cyprus, France, Germany, Greece, Ireland, Italy, The Netherlands, Rumania, Spain, U.K)



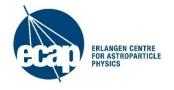
TDR published: http://www.km3net.org/TDR/KM3NeT-TDR.pdf (ISBN 978-90-6488-033-9)

South Pole and Mediterranean Fields of View

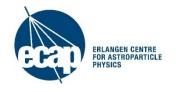


The Objectives

- Central physics goals:
 - Investigate neutrino "point sources" in energy regime 1-100 TeV
 - Complement IceCube field of view
 - Substantially exceed IceCube sensitivity
 - Topics not used for optimisation:
 - Dark Matter
 - Neutrino particle physics aspects
 - Exotics (Magnetic Monopoles, Lorentz invariance violation, ...)
- Implementation requirements:
 - Construction time ≤5 years
 - Operation over at least 10 years without "major maintenance"

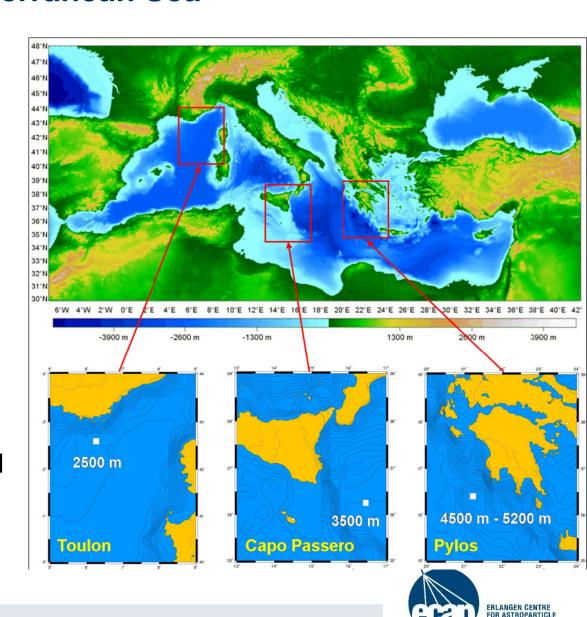


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KM3NeT in the Mediterranean Sea

- characterisation
 measurements
 performed during the
 Design Study at three
 different locations:
 Toulon (ANTARES),
 Capo Passero (NEMO)
 Pylos area (NESTOR)
- Infrastructure of networked KM3NeT nodes foreseen



Technical Design

Objective: Support 3D-array of photodetectors and connect them to shore (data, power, slow control)

- Optical Modules
- Front-end electronics
- Readout, data acquisition, data transport
- — Mechanical structures, backbone cable
- – General deployment strategy
- Sea-bed network: cables, junction boxes
- Calibration devices
 - Shore infrastructure
 - Assembly, transport, logistics
 - Risk analysis and quality control

Unique or preferred solutions

Design rationale:

Cost-effective
Reliable
Producible
Easy to deploy



The Deep Sea environment

One of the most adverse places for a technical installation:

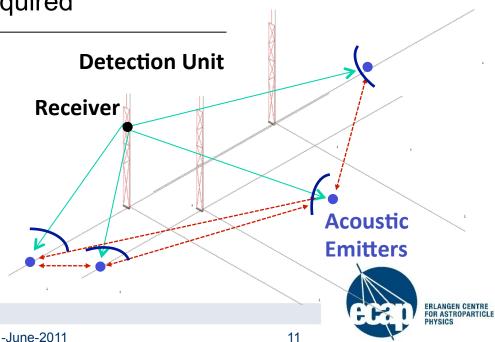
- High pressure (~200bar at 2km)
- Salt water: highly corrosive environment
- Long distance from shore for communication
- Forces on structure due to sea currents

Wet-mateable connectors required

Acoustic position calibration:

Optical calibration:

See poster #13 by U. Emanuele



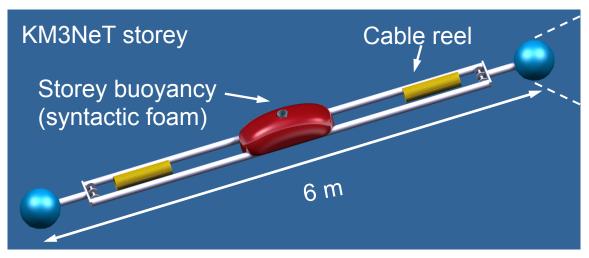


KM3Net Detection Unit

Detection unit (DU): 20 storeys / 40 m distance (DU height ~ 900 m)

storey: bar of 6 m with 2 DOMs

DOM: digital optical module with 31 3" PMT



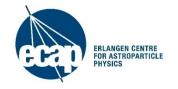


KM3NeT DOM

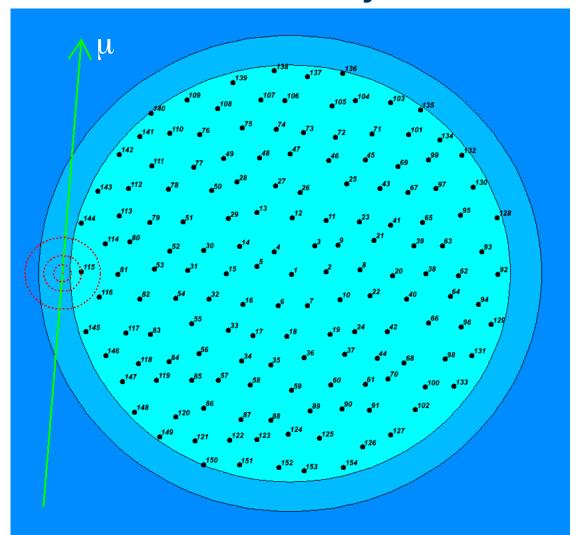
KM3NeT "Building Block" configuration

DU	St/DU	DOM/St.	PMT/DOM	PMT
154	20	2	31	190 960





Detector Geometry



Building Block: Currently considered option: Randomised hexagonial grid

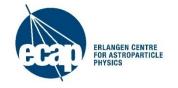
Surface area = π R² = 4.2 km²

R = 1160 m

Instrumented volume = $\pi R^2 h$ h = 760m (19x40) $V_{inst} \approx 3 \text{ km}^3$

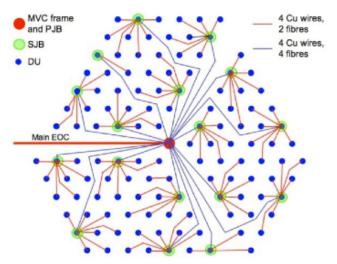
2 Building Blocks will make up KM3NeT

Budget ~220 M€



Data Network and Data Transmission

- All data to shore concept (no trigger undersea)
- Data transport on optical fibers (data, slow control)
 - Optical point-to-point connection DOM-shore ⇒ large number of channels
 - DWDM (Dense Wavelength Division Multiplexing) technique: signals carried by different frequencies (colors) over the same fibre ⇒minimize number of fibers



EOC = electro-optical cable

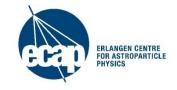
MVC = main voltage converter

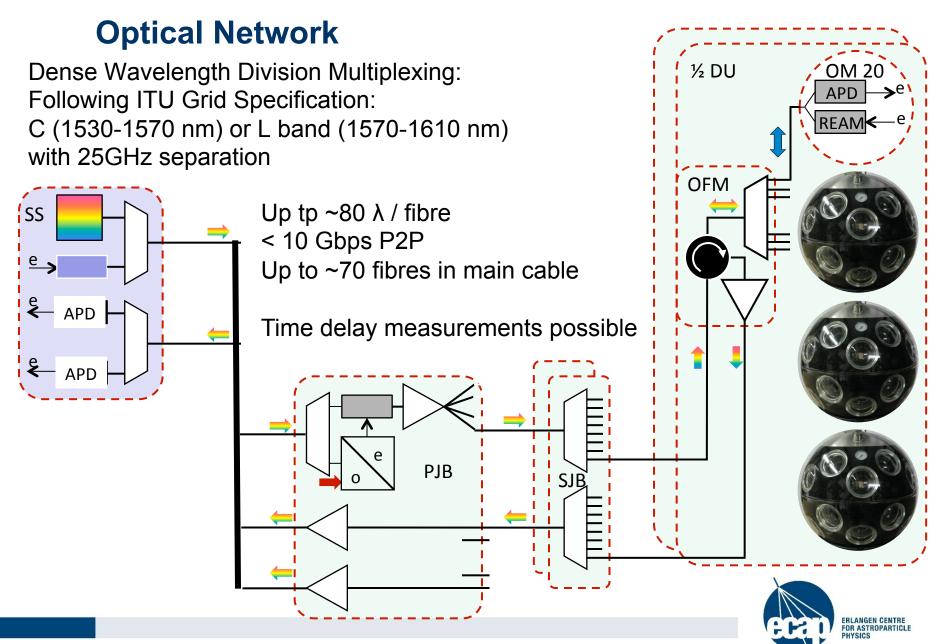
PJB = primary junction box

SJB = secondary junction box

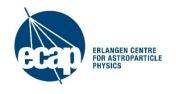
Alternative option: Ring geometry

Star geometry of power and fibre distribution



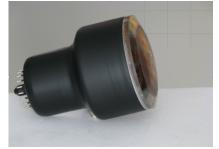


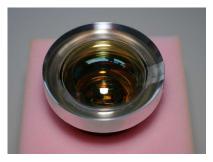
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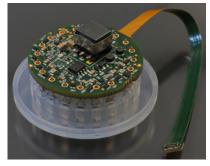
DOM (Digital Optical Module)











Multi-PMT Design:

- Large photocathode area per OM
- Single vs. multi-photon hit separation
- Sphere Ø17"
- PMTs (19+12) ∅3"
- Base
 - Adjust. HV (800-1400V)
 - Comparator for time-over-threshold
- Power board, electronics
- Calibration devices
- Single penetrator



Photo Multiplier Tubes

Requirements:

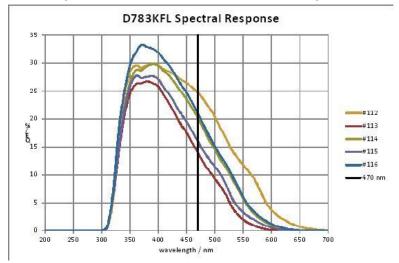
Quantum efficiency (QE) at 404nm Transit time spread (TTS) Gain

Requirements adapted to in-situ conditions:

- Cherenkov photon spectrum in Mediterranean Sea most intense between ~400nm and ~500nm
- Chromatic dispersion sets lower level of required TTS

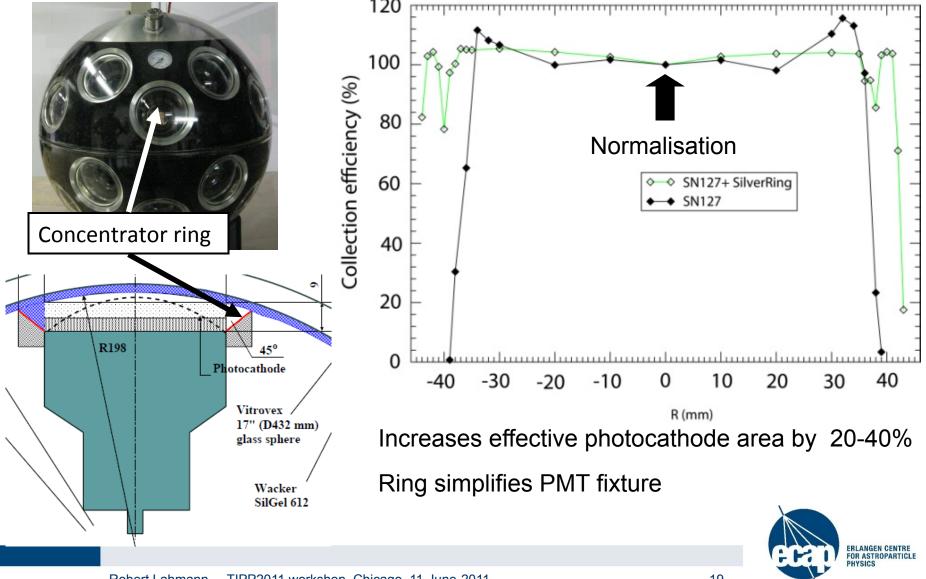
>32% (>20% @ 470nm) <2ns (sigma) 5x10⁶

Prototype tubes from company ETEL:

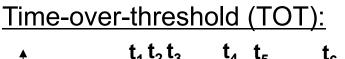


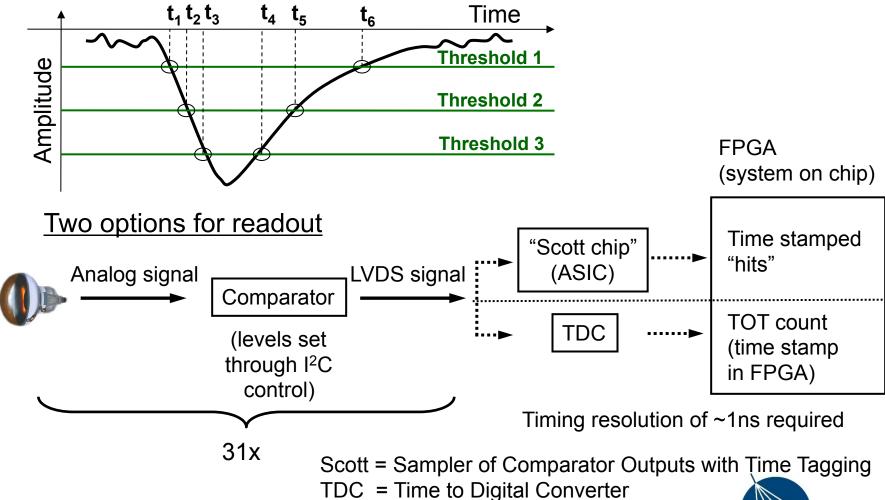


DOMs: Recent Technical Progress



Front End Electronics





Summary and Conclusions

- Major technical design decisions taken, minor points optimised for mass production
- In-situ operation of prototype Detection Unit planned for first half of 2012
- Footprint being optimised for detection of Galactic sources
- Infrastructure of networked KM3NeT nodes most likely scenario
- Data taking could start 2014

Funding provided by EU through FP6 contract no. 011937 and FP7 grant agreement no. 212252

